

# Freudenthal Duality in Gravity: from Groups of Type $E_7$ to Pre-Homogeneous Spaces

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## ABSTRACT

Freudenthal duality can be defined as an anti-involutive, non-linear map acting on symplectic spaces. It was introduced in four-dimensional Maxwell-Einstein theories coupled to a non-linear sigma model of scalar fields.

In this short review, I will consider its relation to the  $U$ -duality Lie groups of type  $E_7$  in extended supergravity theories, and comment on the relation between the Hessian of the black hole entropy and the pseudo-Euclidean, rigid special (pseudo)Kähler metric of the pre-homogeneous spaces associated to the  $U$ -orbits.

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# 1 Freudenthal Duality

We start and consider the following Lagrangian density in four dimensions (*cfr. e.g.* [1]):

$$\mathcal{L} = -\frac{R}{2} + \frac{1}{2}g_{ij}(\varphi)\partial_\mu\varphi^i\partial^\mu\varphi^j + \frac{1}{4}I_{\Lambda\Sigma}(\varphi)F_{\mu\nu}^\Lambda F^{\Sigma|\mu\nu} + \frac{1}{8\sqrt{-G}}R_{\Lambda\Sigma}(\varphi)\epsilon^{\mu\nu\rho\sigma}F_{\mu\nu}^\Lambda F_{\rho\sigma}^\Sigma, \quad (1)$$

describing Einstein gravity coupled to Maxwell (Abelian) vector fields and to a non-linear sigma model of scalar fields (with no potential); note that  $\mathcal{L}$  may -but does not necessarily need to - be conceived as the bosonic sector of  $D = 4$  (*ungauged*) supergravity theory. Out of the Abelian two-form field strengths  $F^\Lambda$ 's, one can define their duals  $G_\Lambda$ , and construct a symplectic vector :

$$H := (F^\Lambda, G_\Lambda)^T, \quad {}^*G_{\Lambda|\mu\nu} := 2\frac{\delta\mathcal{L}}{\delta F^{\Lambda|\mu\nu}}. \quad (2)$$

We then consider the simplest solution of the equations of motion deriving from  $\mathcal{L}$ , namely a static, spherically symmetric, asymptotically flat, dyonic extremal black hole with metric [2]

$$ds^2 = -e^{2U(\tau)}dt^2 + e^{-2U(\tau)}\left[\frac{d\tau^2}{\tau^4} + \frac{1}{\tau^2}(d\theta^2 + \sin\theta d\psi^2)\right], \quad (3)$$

where  $\tau := -1/r$ . Thus, the two-form field strengths and their duals can be fluxed on the two-sphere at infinity  $S_\infty^2$  in such a background, respectively yielding the electric and magnetic charges of the black hole itself, which can be arranged in a symplectic vector  $\mathcal{Q}$  :

$$p^\Lambda := \frac{1}{4\pi}\int_{S_\infty^2} F^\Lambda, \quad q_\Lambda := \frac{1}{4\pi}\int_{S_\infty^2} G_\Lambda, \quad (4)$$

$$\mathcal{Q} := (p^\Lambda, q_\Lambda)^T. \quad (5)$$

Then, by exploiting the symmetries of the background (3), the Lagrangian (1) can be dimensionally reduced from  $D = 4$  to  $D = 1$ , obtaining a 1-dimensional effective Lagrangian ( $' := d/d\tau$ ) [3]:

$$\mathcal{L}_{D=1} = (U')^2 + g_{ij}(\varphi)\varphi^{i'}\varphi^{j'} + e^{2U}V_{BH}(\varphi, \mathcal{Q}) \quad (6)$$

along with the Hamiltonian constraint [3]

$$(U')^2 + g_{ij}(\varphi)\varphi^{i'}\varphi^{j'} - e^{2U}V_{BH}(\varphi, \mathcal{Q}) = 0. \quad (7)$$

The so-called “effective black hole potential”  $V_{BH}$  appearing in (6) and (7) is defined as [3]

$$V_{BH}(\varphi, \mathcal{Q}) := -\frac{1}{2}\mathcal{Q}^T \mathcal{M}(\varphi) \mathcal{Q}, \quad (8)$$

in terms of the symplectic and symmetric matrix [1]

$$\mathcal{M} := \begin{pmatrix} \mathbb{I} & -R \\ 0 & \mathbb{I} \end{pmatrix} \begin{pmatrix} I & 0 \\ 0 & I^{-1} \end{pmatrix} \begin{pmatrix} \mathbb{I} & 0 \\ -R & \mathbb{I} \end{pmatrix} = \begin{pmatrix} I + RI^{-1}R & -RI^{-1} \\ -I^{-1}R & I^{-1} \end{pmatrix}, \quad (9)$$

$$\mathcal{M}^T = \mathcal{M}; \quad \mathcal{M}\Omega\mathcal{M} = \Omega, \quad (10)$$

where  $\mathbb{I}$  denotes the identity, and  $R(\varphi)$  and  $I(\varphi)$  are the scalar-dependent matrices occurring in (1); moreover,  $\Omega$  stands for the symplectic metric ( $\Omega^2 = -\mathbb{I}$ ). Note that, regardless of the invertibility of  $R(\varphi)$  and as a consequence of the physical consistence of the kinetic vector matrix  $I(\varphi)$ ,  $\mathcal{M}$  is negative-definite; thus, the effective black hole potential (8) is positive-definite.

By virtue of the matrix  $\mathcal{M}$ , one can introduce a (scalar-dependent) *anti-involution*  $\mathcal{S}$  in any Maxwell-Einstein-scalar theory described by (1) with a symplectic structure  $\Omega$ , as follows :

$$\mathcal{S}(\varphi) : = \Omega \mathcal{M}(\varphi); \quad (11)$$

$$\mathcal{S}^2(\varphi) = \Omega \mathcal{M}(\varphi) \Omega \mathcal{M}(\varphi) = \Omega^2 = -\mathbb{I}; \quad (12)$$

in turn, this allows to define an anti-involution on the dyonic charge vector  $\mathcal{Q}$ , which has been called (scalar-dependent) *Freudenthal duality* [4, 5, 6] :

$$\mathfrak{F}(\mathcal{Q}; \varphi) : = -\mathcal{S}(\varphi) \mathcal{Q}; \quad (13)$$

$$\mathfrak{F}^2 = -\mathbb{I}, \quad (\forall \{\varphi\}). \quad (14)$$

By recalling (8) and (11), the action of  $\mathfrak{F}$  on  $\mathcal{Q}$ , defining the so-called ( $\varphi$ -dependent) Freudenthal dual of  $\mathcal{Q}$  itself, can be related to the symplectic gradient of the effective black hole potential  $V_{BH}$  :

$$\mathfrak{F}(\mathcal{Q}; \varphi) = \Omega \frac{\partial V_{BH}(\varphi, \mathcal{Q})}{\partial \mathcal{Q}}. \quad (15)$$

Through the attractor mechanism [7], all this enjoys an interesting physical interpretation when evaluated at the (unique) event horizon of the extremal black hole (3) (denoted below by the subscript “ $H$ ”); indeed

$$\partial_\varphi V_{BH} = 0 \Leftrightarrow \lim_{\tau \rightarrow -\infty} \varphi^i(\tau) = \varphi_H^i(\mathcal{Q}); \quad (16)$$

$$S_{BH}(\mathcal{Q}) = \frac{A_H}{4} = \pi V_{BH}|_{\partial_\varphi V_{BH}=0} = -\frac{\pi}{2} \mathcal{Q}^T \mathcal{M}_H(\mathcal{Q}) \mathcal{Q}, \quad (17)$$

where  $S_{BH}$  and  $A_H$  respectively denote the Bekenstein-Hawking entropy [8] and the area of the horizon of the extremal black hole, and the matrix horizon value  $\mathcal{M}_H$  is defined as

$$\mathcal{M}_H(\mathcal{Q}) := \lim_{\tau \rightarrow -\infty} \mathcal{M}(\varphi(\tau)). \quad (18)$$

Correspondingly, one can define the (scalar-independent) horizon Freudenthal duality  $\mathfrak{F}_H$  as the horizon limit of (13) :

$$\tilde{\mathcal{Q}} \equiv \mathfrak{F}_H(\mathcal{Q}) := \lim_{\tau \rightarrow -\infty} \mathfrak{F}(\mathcal{Q}; \varphi(\tau)) = -\Omega \mathcal{M}_H(\mathcal{Q}) \mathcal{Q} = \frac{1}{\pi} \Omega \frac{\partial S_{BH}(\mathcal{Q})}{\partial \mathcal{Q}}. \quad (19)$$

Remarkably, the (horizon) Freudenthal dual of  $\mathcal{Q}$  is nothing but ( $1/\pi$  times) the symplectic gradient of the Bekenstein-Hawking black hole entropy  $S_{BH}$ ; this latter, from dimensional considerations, is only constrained to be an homogeneous function of degree two in  $\mathcal{Q}$ . As a result,  $\tilde{\mathcal{Q}} = \tilde{\mathcal{Q}}(\mathcal{Q})$  is generally a complicated (non-linear) function, homogeneous of degree one in  $\mathcal{Q}$ .

It can be proved that the entropy  $S_{BH}$  itself is invariant along the flow in the charge space  $\mathcal{Q}$  defined by the symplectic gradient (or, equivalently, by the horizon Freudenthal dual) of  $\mathcal{Q}$  itself :

$$S_{BH}(\mathcal{Q}) = S_{BH}(\mathfrak{F}_H(\mathcal{Q})) = S_{BH}\left(\frac{1}{\pi} \Omega \frac{\partial S_{BH}(\mathcal{Q})}{\partial \mathcal{Q}}\right) = S_{BH}(\tilde{\mathcal{Q}}). \quad (20)$$

It is here worth pointing out that this invariance is pretty remarkable : the (semi-classical) Bekenstein-Hawking entropy of an extremal black hole turns out to be invariant under a generally non-linear map acting on the black hole charges themselves, and corresponding to a symplectic gradient flow in their corresponding vector space.

For other applications and instances of Freudenthal duality, see [9, 10, 11].

## 2 Groups of Type $E_7$

The concept of Lie groups *of type*  $E_7$  as introduced in the 60s by Brown [12], and then later developed *e.g.* by [13, 14, 15, 16, 17].

Starting from a pair  $(G, \mathbf{R})$  made of a Lie group  $G$  and its faithful representation  $\mathbf{R}$ , the three axioms defining  $(G, \mathbf{R})$  as a group of type  $E_7$  read as follows :

1. Existence of a (unique) symplectic invariant structure  $\Omega$  in  $\mathbf{R}$  :

$$\exists! \Omega \equiv \mathbf{1} \in \mathbf{R} \times_a \mathbf{R}, \quad (21)$$

which then allows to define a symplectic product  $\langle \cdot, \cdot \rangle$  among two vectors in the representation space  $\mathbf{R}$  itself :

$$\langle Q_1, Q_2 \rangle := Q_1^M Q_2^N \Omega_{MN} = - \langle Q_2, Q_1 \rangle. \quad (22)$$

2. Existence of (unique) rank-4 completely symmetric invariant tensor ( $K$ -tensor) in  $\mathbf{R}$  :

$$\exists! K \equiv \mathbf{1} \in (\mathbf{R} \times \mathbf{R} \times \mathbf{R} \times \mathbf{R})_s, \quad (23)$$

which then allows to define a degree-4 invariant polynomial  $I_4$  in  $\mathbf{R}$  itself :

$$I_4 := K_{MNPQ} Q^M Q^N Q^P Q^Q. \quad (24)$$

3. Defining a triple map  $T$  in  $\mathbf{R}$  as

$$T : \mathbf{R} \times \mathbf{R} \times \mathbf{R} \rightarrow \mathbf{R}; \quad (25)$$

$$\langle T(Q_1, Q_2, Q_3), Q_4 \rangle := K_{MNPQ} Q_1^M Q_2^N Q_3^P Q_4^Q, \quad (26)$$

it holds that

$$\langle T(Q_1, Q_1, Q_2), T(Q_2, Q_2, Q_2) \rangle = \langle Q_1, Q_2 \rangle K_{MNPQ} Q_1^M Q_2^N Q_2^P Q_2^Q. \quad (27)$$

This property makes a group of type  $E_7$  amenable to a description as an automorphism group of a *Freudenthal triple system* (or, equivalently, as the conformal groups of the underlying Jordan triple system - whose a Jordan algebra is a particular case - ).

All electric-magnetic duality ( $U$ -duality<sup>1</sup>) groups of  $\mathcal{N} \geq 2$ -extended  $D = 4$  supergravity theories with symmetric scalar manifolds are of type  $E_7$ . Among these, degenerate groups of type  $E_7$  are those in which the  $K$ -tensor is actually reducible, and thus  $I_4$  is the square of a quadratic invariant polynomial  $I_2$ . In fact, in general, in theories with electric-magnetic duality groups of type  $E_7$  holds that

$$S_{BH} = \pi \sqrt{|I_4(\mathcal{Q})|} = \pi \sqrt{|K_{MNPQ} \mathcal{Q}^M \mathcal{Q}^N \mathcal{Q}^P \mathcal{Q}^Q|}, \quad (28)$$

whereas in the case of degenerate groups of type  $E_7$  it holds that  $I_4(\mathcal{Q}) = (I_2(\mathcal{Q}))^2$ , and therefore the latter formula simplifies to

$$S_{BH} = \pi \sqrt{|I_4(\mathcal{Q})|} = \pi |I_2(\mathcal{Q})|. \quad (29)$$

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<sup>1</sup>Here  $U$ -duality is referred to as the “continuous” symmetries of [18]. Their discrete versions are the  $U$ -duality non-perturbative string theory symmetries introduced by Hull and Townsend [19].

$J_3$	$G_4$	<b>R</b>	$\mathcal{N}$
$J_3^{\mathbb{O}}$	$E_{7(-25)}$	<b>56</b>	2
$J_3^{\mathbb{O}_s}$	$E_{7(7)}$	<b>56</b>	8
$J_3^{\mathbb{H}}$	$SO^*(12)$	<b>32</b>	2, 6
$J_3^{\mathbb{H}_s}$	$SO(6, 6)$	<b>32</b>	0
$J_3^{\mathbb{C}}$	$SU(3, 3)$	<b>20</b>	2
$J_3^{\mathbb{C}_s}$	$SL(6, \mathbb{R})$	<b>20</b>	0
$M_{1,2}(\mathbb{O})$	$SU(1, 5)$	<b>20</b>	5
$J_3^{\mathbb{R}}$	$Sp(6, \mathbb{R})$	<b>14'</b>	2
$\mathbb{R} \oplus \mathbb{R} \oplus \mathbb{R}$ (STU)	$[SL(2, \mathbb{R})]^3$	<b>(2, 2, 2)</b>	2
$\mathbb{R}$ ( $T^3$ )	$SL(2, \mathbb{R})$	<b>4</b>	2

Table 1: Simple, non-degenerate groups  $G$  related to Freudenthal triple systems  $\mathfrak{M}(J_3)$  on simple rank-3 Jordan algebras  $J_3$ . In general,  $G \cong \text{Conf}(J_3) \cong \text{Aut}(\mathfrak{M}(J_3))$  (see *e.g.* [20, 21, 22] for a recent introduction, and a list of Refs.).  $\mathbb{O}$ ,  $\mathbb{H}$ ,  $\mathbb{C}$  and  $\mathbb{R}$  respectively denote the four division algebras of octonions, quaternions, complex and real numbers, and  $\mathbb{O}_s$ ,  $\mathbb{H}_s$ ,  $\mathbb{C}_s$  are the corresponding split forms. Note that the  $G$  related to split forms  $\mathbb{O}_s$ ,  $\mathbb{H}_s$ ,  $\mathbb{C}_s$  is the *maximally non-compact (split)* real form of the corresponding compact Lie group.  $M_{1,2}(\mathbb{O})$  is the Jordan triple system generated by  $2 \times 1$  vectors over  $\mathbb{O}$  [23]. Note that the STU model, based on  $J_3 = \mathbb{R} \oplus \mathbb{R} \oplus \mathbb{R}$ , has a *semi-simple*  $G_4$ , but its *triality symmetry* [24] renders it “effectively simple”. The  $D = 5$  uplift of the  $T^3$  model based on  $J_3 = \mathbb{R}$  is the *pure*  $\mathcal{N} = 2$ ,  $D = 5$  supergravity.  $J_3^{\mathbb{H}}$  is related to both 8 and 24 supersymmetries, because the corresponding supergravity theories are “*twin*”, namely they share the very same bosonic sector [23, 25, 26, 27].

Simple, non-degenerate groups of type  $E_7$  relevant to  $\mathcal{N} \geq 2$ -extended  $D = 4$  supergravity theories with symmetric scalar manifolds are reported in Table 1.

Semi-simple, non-degenerate groups of type  $E_7$  of the same kind are given by  $G = SL(2, \mathbb{R}) \times$

$SO(2, n)$  and  $G = SL(2, \mathbb{R}) \times SO(6, n)$ , with  $\mathbf{R} = (\mathbf{2}, \mathbf{2} + \mathbf{n})$  and  $\mathbf{R} = (\mathbf{2}, \mathbf{6} + \mathbf{n})$ , respectively relevant for  $\mathcal{N} = 2$  and  $\mathcal{N} = 4$  supergravity.

Moreover, degenerate (simple) groups of type  $E_7$  relevant to the same class of theories are  $G = U(1, n)$  and  $G = U(3, n)$ , with complex fundamental representations  $\mathbf{R} = \mathbf{n} + \mathbf{1}$  and  $\mathbf{R} = \mathbf{3} + \mathbf{n}$ , respectively relevant for  $\mathcal{N} = 2$  and  $\mathcal{N} = 3$  supergravity [16].

The classification of groups of type  $E_7$  is still an open problem, even if some progress have been recently made *e.g.* in [28] (in particular, *cfr.* Table D therein).

In all the aforementioned cases, the scalar manifold is a *symmetric* cosets  $\frac{G}{H}$ , where  $H$  is the maximal compact subgroup (with symmetric embedding) of  $G$ . Moreover, the  $K$ -tensor can generally be expressed as [17]

$$K_{MNPQ} = -\frac{n(2n+1)}{6d} \left[ t_{MN}^\alpha t_{\alpha|PQ} - \frac{d}{n(2n+1)} \Omega_{M(P} \Omega_{Q)N} \right], \quad (30)$$

where  $\dim \mathbf{R} = 2n$  and  $\dim G = d$ , and  $t_{MN}^\alpha$  denotes the symplectic representation of the generators of  $G$  itself. Thus, the horizon Freudenthal duality can be expressed in terms of the  $K$ -tensor as follows [4]:

$$\mathfrak{F}_H(\mathcal{Q})_M \equiv \tilde{\mathcal{Q}}_M = \frac{\partial \sqrt{|I_4(\mathcal{Q})|}}{\partial \mathcal{Q}^M} = \epsilon \frac{2}{\sqrt{|I_4(\mathcal{Q})|}} K_{MNPQ} \mathcal{Q}^N \mathcal{Q}^P \mathcal{Q}^Q, \quad (31)$$

where  $\epsilon := I_4 / |I_4|$ ; note that the horizon Freudenthal dual of a given symplectic dyonic charge vector  $\mathcal{Q}$  is well defined only when  $\mathcal{Q}$  is such that  $I_4(\mathcal{Q}) \neq 0$ . Consequently, the invariance (20) of the black hole entropy under the horizon Freudenthal duality can be recast as the invariance of  $I_4$  itself :

$$I_4(\mathcal{Q}) = I_4(\tilde{\mathcal{Q}}) = I_4 \left( \Omega \frac{\partial \sqrt{|I_4(\mathcal{Q})|}}{\partial \mathcal{Q}} \right). \quad (32)$$

In absence of “flat directions” at the attractor points (namely, of unstabilized scalar fields at the horizon of the black hole), and for  $I_4 > 0$ , the expression of the matrix  $\mathcal{M}_H(\mathcal{Q})$  at the horizon can be computed to read

$$\mathcal{M}_{H|MN}(\mathcal{Q}) = -\frac{1}{\sqrt{I_4}} \left( 2\tilde{\mathcal{Q}}_M \tilde{\mathcal{Q}}_N - 6K_{MNPQ} \mathcal{Q}^P \mathcal{Q}^Q + \mathcal{Q}_M \mathcal{Q}_N \right), \quad (33)$$

and it is invariant under horizon Freudenthal duality :

$$\mathfrak{F}_H(\mathcal{M}_H)_{MN} := \mathcal{M}_{H|MN}(\tilde{\mathcal{Q}}) = \mathcal{M}_{H|MN}(\mathcal{Q}). \quad (34)$$

### 3 Duality Orbits, Rigid Special Kähler Geometry and Pre-Homogeneous Vector Spaces

For  $I_4 > 0$ ,  $\mathcal{M}_H(\mathcal{Q})$  given by (33) is one of the two possible solutions to the set of equations [29]

$$\begin{cases} M^T(\mathcal{Q}) \Omega M(\mathcal{Q}) = \epsilon \Omega; \\ M^T(\mathcal{Q}) = M(\mathcal{Q}); \\ \mathcal{Q}^T M(\mathcal{Q}) \mathcal{Q} = -2\sqrt{|I_4(\mathcal{Q})|}, \end{cases} \quad (35)$$

which describes symmetric, purely  $\mathcal{Q}$ -dependent structures at the horizon; they are symplectic or anti-symplectic, depending on whether  $I_4 > 0$  or  $I_4 < 0$ , respectively. Since in the class

of (super)gravity  $D = 4$  theories discussed the sign of  $I_4$  actually determines a stratification of the representation space  $\mathbf{R}$  of charges into distinct orbits of the action of  $G$  into  $\mathbf{R}$  itself (usually named duality orbits), the symplectic or anti-symplectic nature of the solutions to the system (35) is  $G$ -invariant, and supported by the various duality orbits of  $G$  (in particular, by the so-called “large” orbits, for which  $I_4$  is non-vanishing).

One of the two possible solutions to the system (35) reads [29]

$$M_+(\mathcal{Q}) = -\frac{1}{\sqrt{|I_4|}} \left( 2\tilde{\mathcal{Q}}_M \tilde{\mathcal{Q}}_N - 6\epsilon K_{MNPQ} \mathcal{Q}^P \mathcal{Q}^Q + \epsilon \mathcal{Q}_M \mathcal{Q}_N \right);$$

$$\mathfrak{F}_H (M_+)_{MN} : = M_{+|MN}(\tilde{\mathcal{Q}}) = \epsilon M_{+|MN}(\mathcal{Q}).$$

For  $\epsilon = +1 \Leftrightarrow I_4 > 0$ , it thus follows that

$$M_+(\mathcal{Q}) = \mathcal{M}_H(\mathcal{Q}), \quad (36)$$

as anticipated.

On the other hand, the other solution to system (35) reads [29]

$$M_-(\mathcal{Q}) = \frac{1}{\sqrt{|I_4|}} \left( \tilde{\mathcal{Q}}_M \tilde{\mathcal{Q}}_N - 6\epsilon K_{MNPQ} \mathcal{Q}^P \mathcal{Q}^Q \right); \quad (37)$$

$$\mathfrak{F}_H (M_-)_{MN} : = M_{-|MN}(\tilde{\mathcal{Q}}) = \epsilon M_{-|MN}(\mathcal{Q}). \quad (38)$$

By recalling the definition of  $I_4$  (24), it is then immediate to realize that  $M_-(\mathcal{Q})$  is the (opposite of the) Hessian matrix of  $(1/\pi$  times) the black hole entropy  $S_{BH}$  :

$$M_{-|MN}(\mathcal{Q}) = -\partial_M \partial_N \sqrt{|I_4|} = -\frac{1}{\pi} \partial_M \partial_N S_{BH}. \quad (39)$$

The matrix  $M_-(\mathcal{Q})$  is the (opposite of the) pseudo-Euclidean metric of a non-compact, non-Riemannian rigid special Kähler manifold related to the duality orbit of the black hole electromagnetic charges (to which  $\mathcal{Q}$  belongs), which is an example of pre-homogeneous vector space (PVS) [30]. In turn, the nature of the rigid special manifold may be Kähler or pseudo-Kähler, depending on the existence of a  $U(1)$  or  $SO(1, 1)$  connection<sup>2</sup>.

In order to clarify this statement, let us make two examples within maximal  $\mathcal{N} = 8$ ,  $D = 4$  supergravity. In this theory, the electric-magnetic duality group is  $G = E_{7(7)}$ , and the representation in which the e.m. charges sit is its fundamental  $\mathbf{R} = \mathbf{56}$ . The scalar manifold has rank-7 and it is the real symmetric coset<sup>3</sup>  $G/H = E_{7(7)}/SU(8)$ , with dimension 70.

1. The unique duality orbit determined by the  $G$ -invariant constraint  $I_4 > 0$  is the 55-dimensional non-symmetric coset

$$\mathcal{O}_{I_4 > 0} = \frac{E_{7(7)}}{E_{6(2)}}. \quad (40)$$

By customarily assigning positive (negative) signature to non-compact (compact) generators, the pseudo-Euclidean signature of  $\mathcal{O}_{I_4 > 0}$  is  $(n_+, n_-) = (30, 25)$ . In this case,  $M_-(\mathcal{Q})$  given by (39) is the 56-dimensional metric of the non-compact, non-Riemannian rigid special Kähler non-symmetric manifold

$$\mathbf{O}_{I_4 > 0} = \frac{E_{7(7)}}{E_{6(2)}} \times \mathbb{R}^+, \quad (41)$$

with signature  $(n_+, n_-) = (30, 26)$ , thus with character  $\chi := n_+ - n_- = 4$ . Through a conification procedure (amounting to modding out<sup>4</sup>  $\mathbb{C} \cong SO(2) \times SO(1, 1) \cong U(1) \times \mathbb{R}^+$ ,

<sup>2</sup>For a thorough introduction to special Kähler geometry, see *e.g.* [31].

<sup>3</sup>To be more precise, it is worth mentioning that the actual relevant coset manifold is  $E_{7(7)}/[SU(8)/\mathbb{Z}_2]$ , because spinors transform according to the double cover of the stabilizer of the scalar manifold (see *e.g.* [32, 33], and Refs. therein).

<sup>4</sup>The signature along the  $\mathbb{R}^+$ -direction is negative [29].

$G$	$V$	$n$	isotropy alg.	degree	interpr. $D = 4$
$SL(2, \mathbb{C})$	$S^3 \mathbb{C}^2$	1	0	4	$\mathcal{N} = 2, \mathbb{R} (T^3)$
$SL(6, \mathbb{C})$	$\Lambda^3 \mathbb{C}^6$	1	$\mathfrak{sl}(3, \mathbb{C})^{\oplus 2}$	4	$\mathcal{N} = 2, J_3^{\mathbb{C}}$
					$\mathcal{N} = 0, J_3^{\mathbb{C}_s}$
					$\mathcal{N} = 5, M_{1,2}(\mathbb{O})$
$SL(7, \mathbb{C})$	$\Lambda^3 \mathbb{C}^7$	1	$\mathfrak{g}_2^{\mathbb{C}}$	7	
$SL(8, \mathbb{C})$	$\Lambda^3 \mathbb{C}^8$	1	$\mathfrak{sl}(3, \mathbb{C})$	16	
$SL(3, \mathbb{C})$	$S^2 \mathbb{C}^3$	2	0	6	
$SL(5, \mathbb{C})$	$\Lambda^2 \mathbb{C}^5$	3	$\mathfrak{sl}(2, \mathbb{C})$	5	
		4	0	10	
$SL(6, \mathbb{C})$	$\Lambda^2 \mathbb{C}^6$	2	$\mathfrak{sl}(2, \mathbb{C})^{\oplus 3}$	6	
$SL(3, \mathbb{C})^{\otimes 2}$	$\mathbb{C}^3 \otimes \mathbb{C}^3$	2	$\mathfrak{gl}(1, \mathbb{C})^{\oplus 2}$	6	
$Sp(6, \mathbb{C})$	$\Lambda_0^3 \mathbb{C}^6$	1	$\mathfrak{sl}(3, \mathbb{C})$	4	$\mathcal{N} = 2, J_3^{\mathbb{R}}$
$Spin(7, \mathbb{C})$	$\mathbb{C}^8$	1	$\mathfrak{g}_2^{\mathbb{C}}$	2	
		2	$\mathfrak{sl}(3, \mathbb{C}) \oplus \mathfrak{so}(2, \mathbb{C})$	2	
		3	$\mathfrak{sl}(2, \mathbb{C}) \oplus \mathfrak{so}(3, \mathbb{C})$	2	
$Spin(9, \mathbb{C})$	$\mathbb{C}^{16}$	1	$\mathfrak{spin}(7, \mathbb{C})$	2	
$Spin(10, \mathbb{C})$	$\mathbb{C}^{16}$	2	$\mathfrak{g}_2^{\mathbb{C}} \oplus \mathfrak{sl}(2, \mathbb{C})$	2	
		3	$\mathfrak{sl}(2, \mathbb{C}) \oplus \mathfrak{so}(3, \mathbb{C})$	4	
$Spin(11, \mathbb{C})$	$\mathbb{C}^{32}$	1	$\mathfrak{sl}(5, \mathbb{C})$	4	
$Spin(12, \mathbb{C})$	$\mathbb{C}^{32}$	1	$\mathfrak{sl}(6, \mathbb{C})$	4	$\mathcal{N} = 2, 6, J_3^{\mathbb{H}}$
					$\mathcal{N} = 0, J_3^{\mathbb{H}_s}$
$Spin(14, \mathbb{C})$	$\mathbb{C}^{64}$	1	$\mathfrak{g}_2^{\mathbb{C}} \oplus \mathfrak{g}_2^{\mathbb{C}}$	8	
$G_2^{\mathbb{C}}$	$\mathbb{C}^7$	1	$\mathfrak{sl}(3, \mathbb{C})$	2	
		2	$\mathfrak{gl}(2, \mathbb{C})$	2	
$E_6^{\mathbb{C}}$	$\mathbb{C}^{27}$	1	$\mathfrak{f}_4^{\mathbb{C}}$	3	
		2	$\mathfrak{so}(8, \mathbb{C})$	6	
$E_7^{\mathbb{C}}$	$\mathbb{C}^{56}$	1	$\mathfrak{e}_6^{\mathbb{C}}$	4	$\mathcal{N} = 2, J_3^{\mathbb{O}}$
					$\mathcal{N} = 8, J_3^{\mathbb{O}_s}$

Table 2: Non-generic, nor irregular PVS with simple  $G$ , of type 2 (in the complex ground field). To avoid discussing the finite groups appearing, the list presents the Lie algebra of the isotropy group rather than the isotropy group itself [34]. The interpretation (of suitable real, non-compact slices) in  $D = 4$  theories of Einstein gravity is added; remaining cases will be investigated in a forthcoming publication

one can obtain the corresponding 54-dimensional non-compact, non-Riemannian special Kähler symmetric manifold

$$\mathbf{O}_{I_4>0}/\mathbb{C} \cong \widehat{\mathbf{O}}_{I_4>0} = \frac{E_{7(7)}}{E_{6(2)} \times U(1)}. \quad (42)$$

2. The unique duality orbit determined by the  $G$ -invariant constraint  $I_4 < 0$  is the 55-

dimensional non-symmetric coset

$$\mathcal{O}_{I_4<0} = \frac{E_{7(7)}}{E_{6(6)}}, \quad (43)$$

with pseudo-Euclidean signature given by  $(n_+, n_-) = (28, 27)$ , thus with character  $\chi = 0$ . In this case,  $M_-(\mathcal{Q})$  given by (39) is the 56-dimensional metric of the non-compact, non-Riemannian rigid special pseudo-Kähler non-symmetric manifold

$$\mathbf{O}_{I_4<0} = \frac{E_{7(7)}}{E_{6(6)}} \times \mathbb{R}^+, \quad (44)$$

with signature  $(n_+, n_-) = (28, 28)$ . Through a “pseudo-conification” procedure (amounting to modding out  $\mathbb{C}_s \cong SO(1, 1) \times SO(1, 1) \cong \mathbb{R}^+ \times \mathbb{R}^+$ , one can obtain the corresponding 54-dimensional non-compact, non-Riemannian special pseudo-Kähler symmetric manifold

$$\mathbf{O}_{I_4<0}/\mathbb{C}_s \cong \widehat{\mathbf{O}}_{I_4<0} = \frac{E_{7(7)}}{E_{6(6)} \times SO(1, 1)}. \quad (45)$$

(41) and (44) are non-compact, real forms of  $\frac{E_7}{E_6} \times GL(1)$ , which is the type 29 in the classification of regular, pre-homogeneous vector spaces (PVS) worked out by Sato and Kimura in [34]. From its definition, a PVS is a finite-dimensional vector space  $V$  together with a subgroup  $G$  of  $GL(V)$ , such that  $G$  has an open dense orbit in  $V$ . PVS are subdivided into two types (type 1 and type 2), according to whether there exists an homogeneous polynomial on  $V$  which is invariant under the semi-simple (reductive) part of  $G$  itself. For more details, see *e.g.* [30, 35, 36].

In the case of  $\frac{E_7}{E_6} \times GL(1)$ ,  $V$  is provided by the fundamental representation space  $\mathbf{R} = \mathbf{56}$  of  $G = E_7$ , and there exists a quartic  $E_7$ -invariant polynomial  $I_4$  (24) in the **56**;  $H = E_6$  is the isotropy (stabilizer) group.

Amazingly, simple, non-degenerate groups of type  $E_7$  (relevant to  $D = 4$  Einstein (super)gravities with symmetric scalar manifolds) *almost* saturate the list of irreducible PVS with unique  $G$ -invariant polynomial of degree 4 (*cfr.* Table 2); in particular, the parameter  $n$  characterizing each PVS can be interpreted as the number of centers of the regular solution in the (super)gravity theory with electric-magnetic duality ( $U$ -duality) group given by  $G$ . This topic will be considered in detail in a forthcoming publication.

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